

# Antihydrogen Annihilation Reconstruction with the ALPHA Silicon Detector

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## Abstract

The ALPHA experiment has succeeded in trapping antihydrogen, a major milestone on the road to spectroscopic comparisons of antihydrogen with hydrogen. An annihilation vertex detector, which determines the time and position of antiproton annihilations, has been central to this achievement. This detector, an array of double-sided silicon microstrip detector modules arranged in three con-

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centric cylindrical tiers, is sensitive to the passage of charged particles resulting from antiproton annihilation. This article describes the method used to reconstruct the annihilation location and to distinguish the annihilation signal from the cosmic ray background. Recent experimental results using this detector are outlined.

*Keywords:* antihydrogen, antimatter, event reconstruction, vertexing detector, cosmic ray background suppression

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## 1. Introduction

The ALPHA experiment is engaged in the production and magnetic confinement of antihydrogen atoms [1, 2]. With the apparatus located at the Antiproton Decelerator facility (AD) at CERN [3], the ALPHA collaboration intends to perform precision spectroscopic measurements on trapped antihydrogen [4] as a stringent test of CPT symmetry.

As with several other AD experiments [5–7], ALPHA synthesizes antihydrogen atoms by merging positron and antiproton plasmas [8, 9], which themselves are contained in a Penning-Malmberg charged particle trap (Fig. 1). However, without additional confining potentials, the electrically neutral antihydrogen atoms escape the fields used to confine and manipulate the charged antiparticles. To prevent some of the newly formed antiatoms from traveling to the apparatus walls and annihilating, the ALPHA experiment employs a minimum-B neutral atom trap. This neutral trap consists of an array of superconducting magnets which provides a strong inhomogeneous magnetic field [10]. By exploiting the interaction between the magnetic moments of the atoms and the magnetic field gradient, very low energy ( $< 50 \mu\text{eV}$ ), low-field seeking, antihydrogen atoms have been confined for as long as 1000 s [2].

To detect and locate annihilations, ALPHA has constructed a silicon tracking detector [11, 12]. This detector is similar to that used in the ATHENA antihydrogen experiment [13–15], although the ALPHA instrument does not contain CsI crystals for  $\gamma$ -ray detection, and has three layers of modules to

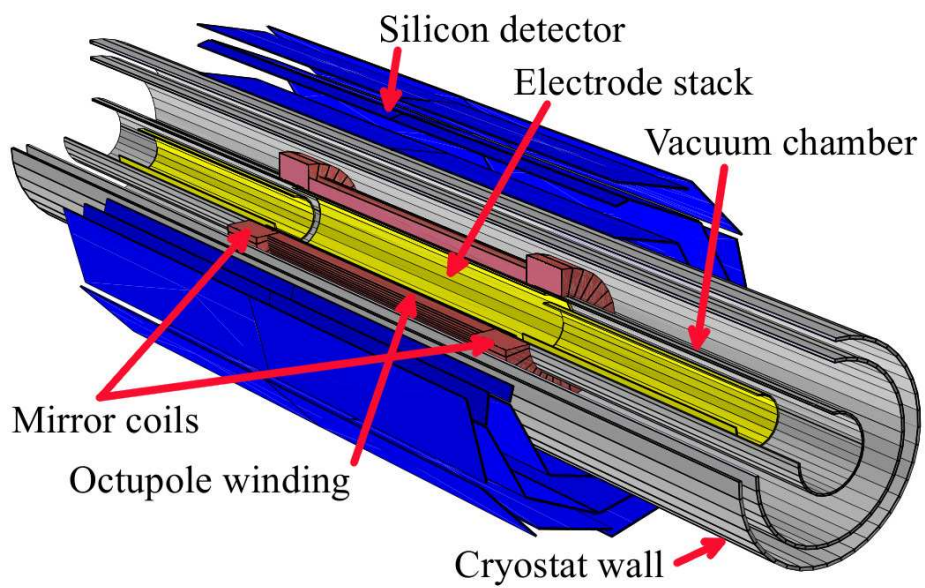


Figure 1: Cut-away diagram of the antihydrogen production and trapping region of the ALPHA apparatus, showing the relative locations of the Penning-Malmberg trap (the external solenoid providing the axial magnetic field is not shown), neutral-atom trap magnets, and silicon detector.

23 ATHENA’s two (see Ref. [16] for design considerations of the ALPHA detec-  
24 tor). The detection method is based on the reconstruction and extrapolation  
25 of the trajectories of charged annihilation products (primarily charged pions),  
26 which enables a 3-dimensional determination of the antiproton annihilation po-  
27 sition, or ‘vertex’.

28 The overarching design consideration for the ALPHA detector was to ensure  
29 compatibility with the rigid experimental requirements necessary for the mag-  
30 netic trapping of antihydrogen. These constraints include the presence of a large  
31 amount of material between the annihilation point and the detector, as well as  
32 limited space available to house the detector. Despite these challenges, the an-  
33 nihilation detection and event reconstruction by this silicon detector provided a  
34 crucial tool for the unambiguous demonstration of antihydrogen trapping [1, 2].  
35 This article reports the details of the vertex reconstruction algorithm, as well  
36 as the analysis method for background suppression, used by the ALPHA exper-  
37 iment. The methods described here are an improved version of those used in  
38 Ref. [1] and were applied to the analysis of the data presented in Ref. [2].

## 39 **2. The ALPHA detector and apparatus**

40 The ALPHA detector (shown in Fig. 2) consists of 60 double-sided silicon  
41 microstrip modules arranged in three concentric layers. The detector is split  
42 axially into two sections, each containing 30 modules. Figure 3 shows the con-  
43 figuration of the silicon modules and their locations with respect to the rest  
44 of the apparatus. The inner and middle layers are situated around the trap  
45 axis with radii of 7.5 cm and 9.55 cm respectively, while the outer layer is split  
46 between radii of 10.9 cm and 11.4 cm.

47 Each detector module has an active silicon area of  $6\text{ cm} \times 23\text{ cm}$ , with 256  
48 readout strips with a pitch width of  $227\text{ }\mu\text{m}$  in the  $R - \phi$  direction, and 256  
49 readout strips with a pitch width of  $875\text{ }\mu\text{m}$  in the  $z$  direction (where  $R$ ,  $\phi$ , and  
50  $z$  are cylindrical coordinates). Since the signal collection strips run in orthog-  
51 onal directions on opposite sides of the silicon wafer, the point of intersection

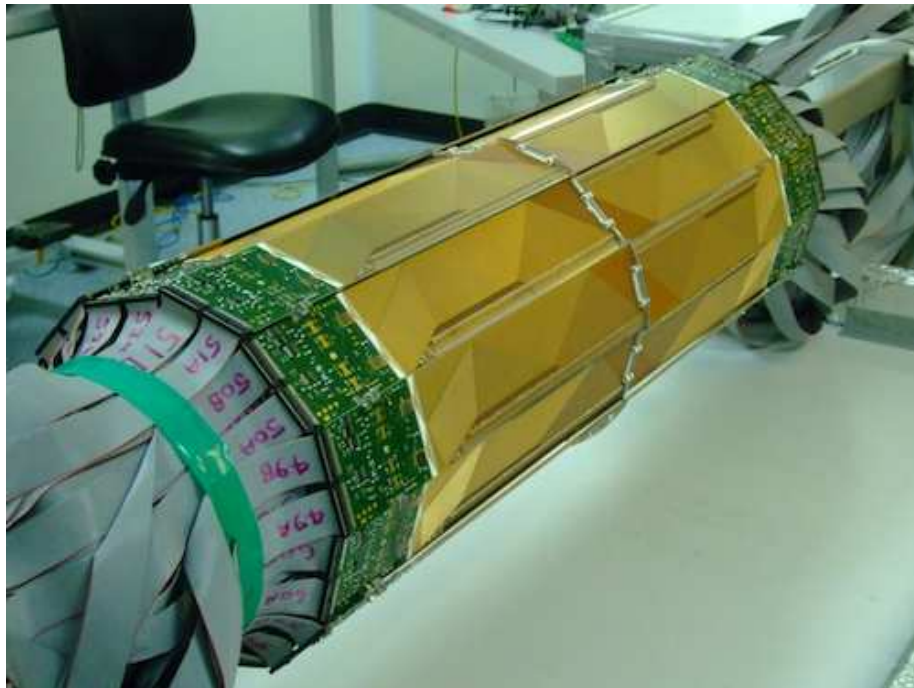


Figure 2: The ALPHA Silicon Detector during construction at the University of Liverpool. The line segments within the gold rectangles are the leads connecting the microstrips to the readout electronics (the silicon wafers and microstrips are located on the opposite side of the modules and cannot be seen). The green rectangular sections contain the on-board readout electronics, and the grey cables carry the analog and digital signals to the rest of the readout system.

52 between the particle trajectory and the silicon module can be localized in the  
53 3-dimensional reference frame of the detector (the point of intersection is nor-  
54 mally called a ‘hit’). The total axial extent of the detector is 46 cm, which  
55 provides a solid angle coverage of  $\sim 90\%$  for annihilations in the axial center.

56 For the purposes of this article, an ‘event’ refers to the full operation of  
57 triggering and digitization of all the signal strips. To coordinate the strip trig-  
58 gering and digitization, every detector module has four VA1TA [17] Application  
59 Specific Integrated Circuits (ASIC), where each ASIC handles 128 strips. The  
60 analog strip signals from the ASIC readout chips are digitized by five 48-channel  
61 VME-based VF48 ADC modules [18]. The programmable trigger condition is  
62 set to read-out the entire detector when two or more  $R - \phi$  strips from the  
63 inner layer of the detector register signal. This trigger is deliberately intended  
64 to accept a broad category of events, as the number of trapped antihydrogen  
65 atoms is known to be small and it is important to accept as many of these rare  
66 events as possible. Through a dedicated cross-calibration with external scin-  
67 tillation detectors with overlapping solid angles, the overall trigger efficiency is  
68 estimated to be  $(90 \pm 10)\%$ .

69 In total, this detector contains 30,720 signal strips, of which 30,195 (or 98.3%)  
70 are fully functioning. The large majority (512 strips) of the absent signal strips  
71 are the result of a non-functioning and disconnected module. The functioning  
72 strips typically operate with a leakage current of  $< 8$  nA per strip at room tem-  
73 perature. The peak read-out rate for this detector is 500 events/s, where each  
74 readout event contains the analog signal output for every strip. The amplitude  
75 of the analog signal is proportional to the number of electron-hole pairs liberated  
76 during the passage of a particle through the silicon volume. A dynamic thresh-  
77 olding algorithm, which utilizes knowledge of the amount of baseline charge  
78 collected for each strip (i.e. the pedestal), is used, on an event-by-event basis,  
79 to determine which strips registered signal. To account for the possibility that  
80 the charge is shared over multiple strips, adjacent strips are grouped together  
81 and the cluster center is determined via the ‘center of gravity’ algorithm [19].  
82 However, because of the large pitch width, the large majority (about 75%) of

83 strips registering signal are unaccompanied by any adjacent signal strips. Thus,  
84 to a good approximation, the hit resolution is  $227/\sqrt{12} = 65 \mu\text{m}$  in the  $R - \phi$   
85 direction, and  $875/\sqrt{12} = 253 \mu\text{m}$  in the  $z$  direction.

86 When reconstructing the paths of charged particles passing through the de-  
87 tector, any materials or fields that might affect the particle trajectories must  
88 be taken into account. In the ALPHA apparatus, annihilation products must  
89 pass through an electrode stack, superconducting magnet windings, and vacuum  
90 chamber walls before reaching the detector (Fig. 3). Moreover, the trap elec-  
91 trodes and silicon detector are surrounded by an external solenoid magnet, which  
92 provides a strong axial magnetic field (typically 1 T). Charged particles with  
93 low transverse kinetic energy will gyrate around the axial magnetic field lines.  
94 Simultaneously, the electrode stack imposes an axial electric field. In addition  
95 to the Penning-Malmberg trap, a significant portion of the ALPHA apparatus  
96 is dedicated to the magnetic neutral trap. In order to provide the maximum  
97 radial field magnitude within the trap region, the superconducting octupolar  
98 magnet is located as close to the trap region as possible [10]. For this reason,  
99 the particles resulting from antiproton annihilation will travel through a large  
100 amount of scattering material before encountering the silicon detector. This  
101 will worsen the resolution of the calculated vertex position, as the reconstructed  
102 trajectories of scattered particles do not lead directly back to the annihilation  
103 position. Specifically, a charged particle traveling outward from the center of  
104 the apparatus will encounter the equivalent of between 40 - 70% of a radiation  
105 length of material depending on the track angle, and whether the trajectory of  
106 the particle encounters the superconducting winding of the neutral atom trap.  
107 This scattering can cause the calculated particle trajectory to deviate from its  
108 actual trajectory by as much as several millimetres for an extrapolated track  
109 with path length of about 5-15 cm.

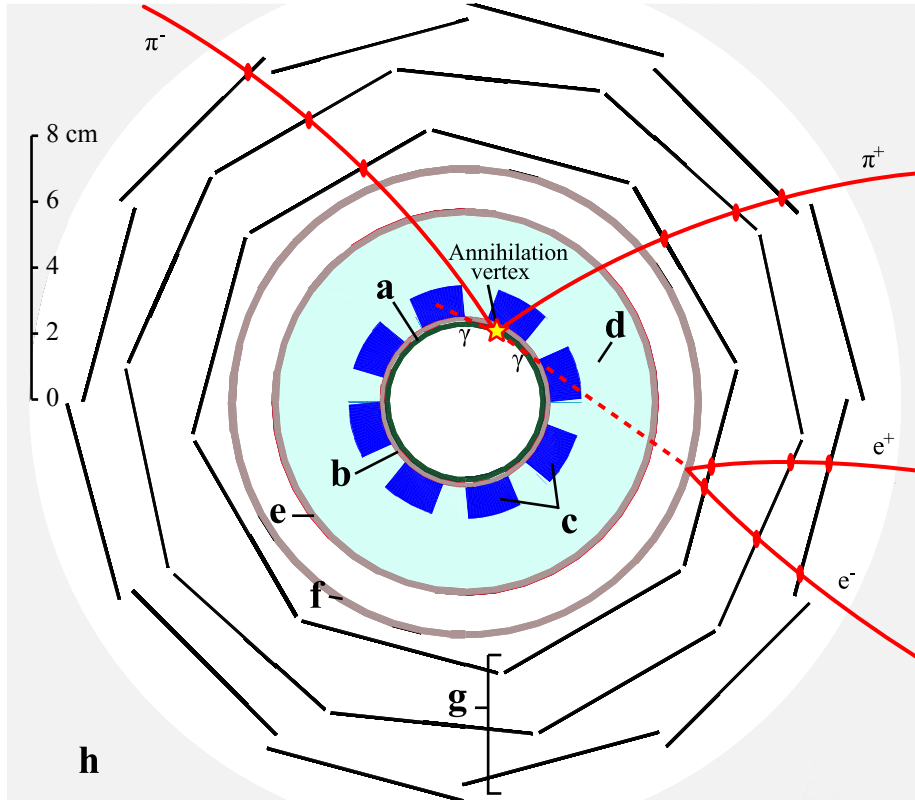


Figure 3: Cross-sectional schematic of the axial center of the ALPHA apparatus and detector (to scale). The labeled apparatus elements are as follows: a) electrode stack, b) magnet winding form, c) octupole magnet winding, d) liquid helium volume, e) inner isolation vacuum wall, f) outer isolation vacuum wall, g) silicon detector, and h) external solenoid magnet. The two mirror coils at the axial ends of the magnetic neutral atom trap are not shown. A cartoon illustration of an antiproton annihilation resulting in three pions (two charged and one neutral) is also shown, where the annihilation vertex is given as the yellow star. The curves represent the trajectories of the annihilation products, with the ovals indicating where the particles passed through a silicon module. The neutral pion has quickly decayed through the  $\pi^0 \rightarrow 2\gamma$  channel. One of the resulting photons has its energy attenuated and is finally absorbed in the octupole winding, while the other photon produced an electron-positron pair.



### 110 **3. Annihilation position reconstruction**

111 The ALPHA silicon detector event reconstruction algorithms attempt to de-  
112 termine the antiproton annihilation position within the ALPHA apparatus. This  
113 task is divided into two parts: 1) the identification and reconstruction of the  
114 trajectories (tracks) of the charged particles released during antiproton annihila-  
115 tion (track reconstruction), and 2) the determination of the primary annihilation  
116 position using the track information (vertex determination). Throughout the  
117 reconstruction process, knowledge of the detector geometry and the outgoing  
118 particles' characteristics is exploited to optimize the overall procedure.

#### 119 *3.1. Track reconstruction*

120 Low-energy antiproton annihilation produces a relatively small number of  
121 particles compared, for example, to a hadron collider environment. On average,  
122 an antiproton annihilation (either on the electrode surface, or on residual gas  
123 atoms) produces about three charged pions and two neutral pions [20]. While  
124 charged pions are stable on the timescale required to reach the detector, neutral  
125 pions decay essentially instantly ( $\sim 10^{-16}$  s) into  $\gamma$ -rays, which, in turn, will  
126 often produce  $e^-e^+$  pairs when transiting the apparatus material (Fig 3). It  
127 should also be noted that stopped antiprotons can also fragment gold nuclei  
128 residing on the electrode surface, resulting in residual product nuclei along with  
129  $\alpha$ ,  $\beta$ , and  $\gamma$ -ray backgrounds [21]. However, due to the amount of apparatus  
130 material, very few massive fragmentation products will arrive at the detector.  
131 Likewise, a large fraction of the  $\gamma$ -ray background from nuclear fragmentation  
132 is attenuated by the material of the neutral-atom trap.

133 In the ALPHA detector, most events will have between 9-15 hits (where a  
134 'hit' is the intersection of orthogonal strips, translated into the 3-dimensional  
135 reference frame of the detector, as discussed in Sec. 2). The small average  
136 number of hits per event allows for the full examination of all hit combinations,  
137 and in turn provides a significant advantage for track finding. For example,  
138 'ghost hit' ambiguities (where two or more particles pass through a single de-  
139 tector module and the orthogonal strip geometry results in several false hits)

140 are resolved in a straightforward way: track candidates containing ghost hits  
141 do not conform to a helical trajectory and do not survive the track selection  
142 criteria. In the same situation, there will also be a track candidate containing  
143 the ‘true’ hits, and this candidate will be much more likely to satisfy the track  
144 selection criteria.

145 The ALPHA detector is located outside the trap cryostat, and therefore  
146 outside of the scattering material and inhomogeneous magnetic field of the  
147 neutral-atom trap, but still within the strong axial magnetic field of the Penning-  
148 Malmberg trap. As such, the trajectories of the charged particles as they pass  
149 through the detector are, to a good approximation, helical. A Monte Carlo  
150 simulation was implemented using GEANT3 within the ROOT Virtual Monte  
151 Carlo software package [15, 22–24] with a realistic magnetic field map generated  
152 with the TOSCA/OPERA3D field solver package [25]. This simulation allowed  
153 for the study of the passage of charged particles through the apparatus mate-  
154 rial and field. Apart from deviations due to multiple scattering, the simulated  
155 charged particles indeed follow helical trajectories in the homogeneous region of  
156 the field containing the silicon detector.

157 Tracks recorded by the ALPHA detector are predominately attributed to  
158 charged pions produced during antiproton annihilation. The detector geome-  
159 try allows each outgoing particle trajectory to be sampled a maximum of three  
160 times, regardless of track angle, which is sufficient to determine the track pa-  
161 rameters needed to extrapolate the helical trajectory in the solenoidal magnetic  
162 field. For this reason, a track candidate in the ALPHA detector is defined as a  
163 collection of exactly three hits (one in each detector layer).

164 In order to identify the charged tracks in an event, every possible three-hit  
165 combination is examined. Each track candidate has six hit degrees of freedom  
166 – that is, each of the three hits has two degrees of freedom in the directions  
167 perpendicular to that of the strips. After determining the five helix parameters,  
168 each track is left with only one effective degree of freedom, which contributes  
169 only in the axial projection of the track. Track candidates can then be selected  
170 based, in part, on how closely the candidates conform to helical trajectories

171 in the axial projection. Specifically, a  $\chi^2$  figure of merit can be constructed  
172 which compares the positions where the determined helix trajectory intersects  
173 the silicon detector modules to the measured hit positions. However, since only  
174 three hits are available for each track candidate, the helix parameters in the  
175 plane perpendicular to the magnetic field are exactly determined. Thus, for  
176 the plane perpendicular to the magnetic field, the computed trajectories always  
177 pass exactly through the hit locations such that these terms do not contribute  
178 to the  $\chi^2$  measure. As such, the  $\chi^2$  figure of merit reflects only how well the  
179 track candidates conform to a helical trajectory in the axial projection, and  
180 only candidates for which  $\chi^2 < 5$  are considered for the vertex reconstruction.  
181 Additionally, track candidates are rejected if their reconstructed trajectories do  
182 not pass within a radius of 3.73 cm from the trap axis (which corresponds to a  
183 volume extending from the trap axis out to a radius 1.5 cm beyond the inner  
184 surface of the Penning-Malmberg trap electrodes).

185 It should also be stressed that since the helix parameters in the plane per-  
186 pendicular to the magnetic field are exactly determined, the track covariance  
187 matrix cannot be fully populated. That is, there is no measure of uncertainty  
188 for the radial trajectory of the track candidates. As a result, without full covari-  
189 ance matrices, sophisticated vertex determination methods (i.e. least-squares  
190 fitting or Kalman filter methods) are precluded.

191 The track finding efficiency of this algorithm is evaluated using the Monte  
192 Carlo simulation described above. After all the track selection criteria are ap-  
193 plied, the track finding efficiency of this algorithm is found to be  $(88 \pm 5)\%$  for  
194 all charged tracks (regardless of particle species) with three hits. This highlights  
195 the advantage of evaluating every hit combination, as only a small number of  
196 tracks are missed due to atypical trajectories through the detector modules,  
197 often because of particle scattering within the silicon itself. However, many of  
198 these properly determined tracks are due to  $e^-e^+$  pairs. As such, although they  
199 return acceptable tracks, these tracks are not guaranteed to extrapolate back to  
200 the annihilation vertex.

201 *3.2. Vertex reconstruction*

202 An annihilation vertex is defined to be a convergence of particle tracks. In  
 203 order to locate such vertices, the particle tracks identified in Section 3.1 are  
 204 extrapolated into the trapping region near the radial center of the ALPHA  
 205 apparatus (Fig. 3). Track extrapolation covers a path length of at least 5.3 cm,  
 206 and as much as  $\sim 14$  cm, where the particle trajectory will necessarily have  
 207 passed through several layers of scattering material. The annihilation vertex is  
 208 taken as the point where the tracks pass closest to each other. The effect of the  
 209 scattering material is to increase the statistical variance of the vertex position  
 210 determination.

211 The measured vertex position,  $\mathbf{r}_{\text{vertex}}$ , is determined through the minimiza-  
 212 tion of a figure of merit,  $D$ , which represents the mean distance of closest ap-  
 213 proach of the tracks to the vertex position:

$$D = \frac{1}{N_{\text{tracks}}} \sum_{i=1}^{N_{\text{tracks}}} d_i, \quad (1)$$

214 where  $N_{\text{tracks}}$  is the number of tracks used in the vertex determination, and  $d_i$   
 215 is the distance of closest approach of the  $i^{\text{th}}$  track, with track position  $\mathbf{r}_i$ , to the  
 216 vertex position:

$$d_i^2 = \min \left\{ |\mathbf{r}_i - \mathbf{r}_{\text{vertex}}|^2 \right\}. \quad (2)$$

217 The minimization of Eq. 1 will then return a vertex position which balances  
 218 the contributions from all the included tracks. Each track is treated equally in  
 219 this procedure, regardless of its dip-angle or the material it encounters along its  
 220 trajectory. As such, it is preferable to include as many tracks as possible, as  
 221 each track will constrain the position of the vertex. However, it is important to  
 222 exclude tracks which do not converge with their counterparts, as these tracks  
 223 can bias the vertex reconstruction away from the true annihilation position.

224 The track exclusion algorithm proceeds as follows:

- 225 1. Do not proceed if  $N_{\text{tracks}} \leq 2$ , as at least two tracks are needed to form  
 226 a vertex.

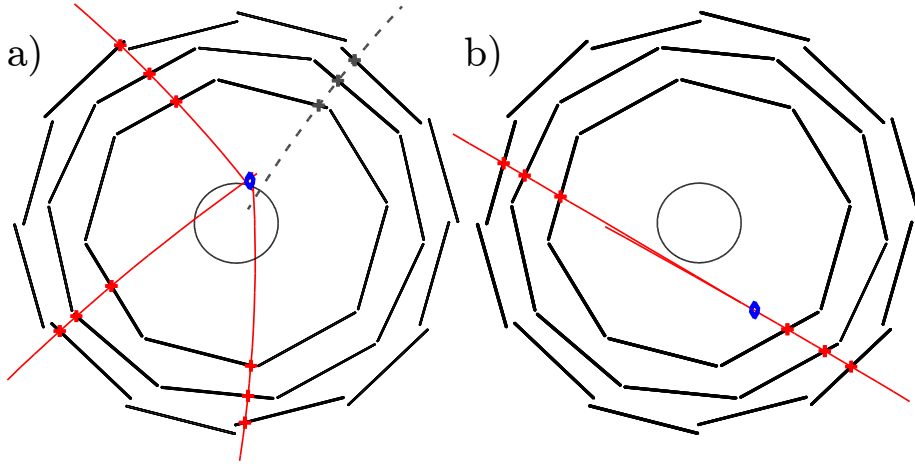


Figure 4: Example reconstruction of a) an annihilation event, and b) a cosmic background event. The crosses indicate hits registered in the silicon microstrip modules, while the solid curves show the reconstructed particle tracks. The dashed track illustrates a track that was considered, but ultimately excluded from the vertex determination. The reconstructed vertex is shown as the hollow diamond.

- 227 2. Reconstruct the vertex position using all the available tracks and calculate
- 228 the mean distance of closest approach for this initial configuration,  $D_0$ .
- 229 Additionally, construct  $N$  auxiliary vertices, where each new vertex con-
- 230 figuration excludes a different track (all of the new vertices will therefore
- 231 include  $N_{\text{tracks}} - 1$  tracks).
- 232 3. Calculate the mean distance of closest approach for each new vertex con-
- 233 figuration. Determine which track configuration has the smallest mean
- 234 distance of closest approach, and call this value  $D_{\text{min}}$ .
- 235 4. Calculate  $\Delta D = (D_0 - D_{\text{min}})/D_0$ , which gives the fractional improvement
- 236 in the mean distance of closest approach by excluding that specific track
- 237 from the vertex.
- 238 5. If  $\Delta D \leq D_{\text{cutoff}}$ , exit the algorithm, keeping the configuration associated
- 239 with  $D_0$  as the final vertex determination.  $D_{\text{cutoff}}$  sets the threshold on the
- 240 fractional change in  $D$ . For the ALPHA detector, a cutoff of  $D_{\text{cutoff}} = 0.4$
- 241 was determined by optimizing the effect of this threshold on the vertex

- 242 resolution using results from the Monte Carlo simulation of our system.
- 243 6. If  $\Delta D > D_{\text{cutoff}}$ , the configuration of tracks associated with  $D_{\text{min}}$  is
- 244 promoted to the current accepted configuration, and relabeled as  $D_0$  for
- 245 the remainder of the algorithm.
- 246 7. If the track configuration now associated with  $D_0$  has more than two
- 247 tracks, return to Step 2 with this configuration.

248 Figure 4 a) shows an example event after the entire reconstruction proce-

249 dure, including track exclusion. In this example, four tracks were identified,

250 but the three solid curves show the tracks that returned the best determination

251 of the vertex position. The resolution of the vertex determination can then be

252 estimated using the Monte Carlo simulation of the antiproton annihilation on

253 the electrode surface. Because the simulated annihilation position is known, the

254 reconstructed position uncertainty can be evaluated, as shown in Fig. 5. Here,

255 the distribution of differences between the simulated and reconstructed posi-

256 tions is fitted with a resolution function comprised of two independent Gaussian

257 terms, whose fitting parameters are given as Table 1. With this resolution func-

258 tion, the narrow Gaussian term represents well-reconstructed vertices, while the

259 broad Gaussian term includes events where the vertex is poorly reconstructed,

260 usually resulting from the inclusion of inappropriate, or inadequately measured,

261 tracks. It is also useful to consider an effective resolution which characterizes

262 the full distribution. As such, the full width at half maximum (FWHM) is

263 used to provide a measure of the overall uncertainty in the reconstructed vertex

264 position. The effective axial resolution is found to be 0.56 cm, while the resolu-

265 tion in the radial component of the vertex position is found to be 0.87 cm, and

266 the azimuthal resolution is  $21.4^\circ$  (which, at the electrode radius of 2.2275 cm,

267 corresponds to 0.83 cm FWHM).

#### 268 4. Cosmic background rejection

269 In addition to charged annihilation products, the ALPHA detector is also

270 sensitive to the passage of charged particles originating from cosmic rays (Fig.

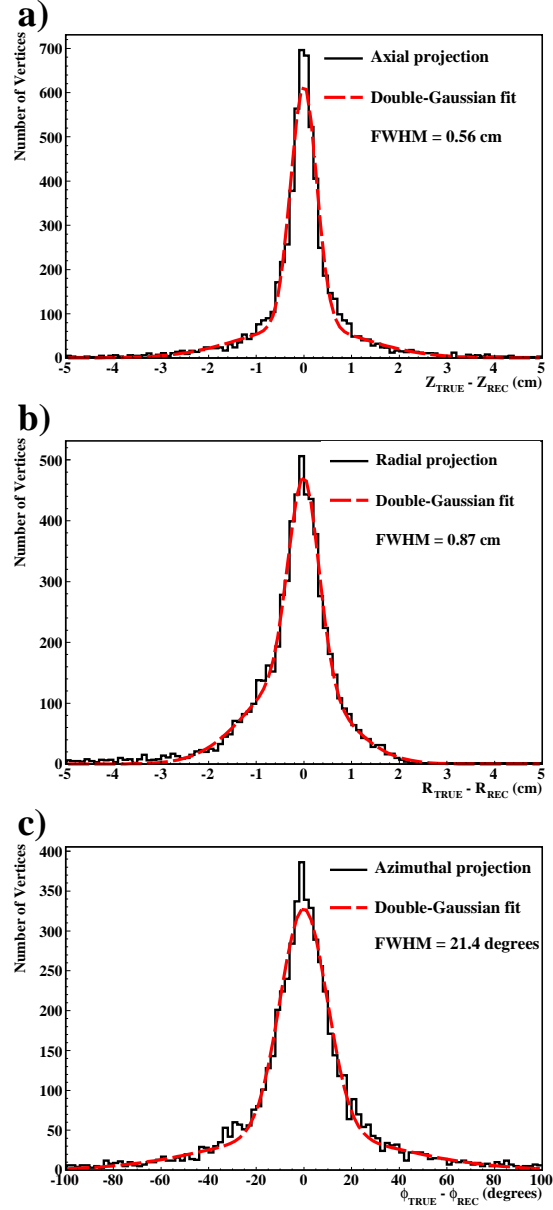


Figure 5: The distributions of differences between true and reconstructed positions for simulated annihilation vertices for the a) axial, b) radial, and c) azimuthal coordinates. The dashed line shows the result of fitting a function with two independent Gaussian terms (denoted broad and narrow), with the fitting parameters given in Table 1.

Coordinate	Axial	Radial	Azimuthal
Narrow Gaussian term:			
Area	$373 \pm 10$	$268 \pm 17$	$7050 \pm 210$
Mean	$-0.004 \pm 0.005$ cm	$-0.004 \pm 0.010$ cm	$-0.11 \pm 0.21^\circ$
Width	$0.27 \pm 0.01$ cm	$0.32 \pm 0.02$ cm	$9.9 \pm 0.3^\circ$
Broad Gaussian term:			
Area	$228 \pm 9$	$344 \pm 17$	$4240 \pm 190$
Mean	$-0.04 \pm 0.03$ cm	$-0.22 \pm 0.02$ cm	$0.52 \pm 0.93^\circ$
Width	$1.29 \pm 0.05$ cm	$0.99 \pm 0.03$ cm	$39.4 \pm 1.5^\circ$

Table 1: Table of fitting parameters for the double-Gaussian resolution functions shown in Figure 5.

271 4b) shows an example cosmic ray event). It is important to reduce the cosmic ray  
272 background in order to assist the identification of antihydrogen annihilations,  
273 especially for the observation of magnetically trapped antihydrogen. The large  
274 majority of cosmic background events are identifiably different from annihilation  
275 events and can be rejected event-by-event.

#### 276 4.1. Discriminating variables

277 Cosmic rays that graze the detector do not typically produce a vertex and  
278 are automatically rejected. A vertex is often reconstructed, however, when the  
279 cosmic ray particle passes through the center of the detector. Fortunately, the  
280 distinct topologies of annihilation and cosmic ray events can be used to classify  
281 events as signal (annihilation, e.g. Fig. 4a) or background (cosmics, e.g. Fig.  
282 4b). There are several variables that can be used to quantify the different signal  
283 and background topology: the number of charged tracks,  $N_{\text{tracks}}$ ; the combined  
284 linear fit residual,  $\delta$ ; and the vertex radial position,  $R$ .

##### 285 4.1.1. Number of charged tracks, $N_{\text{tracks}}$

286 The majority of cosmic background events which return a vertex position  
287 contain two charged tracks ( $N_{\text{tracks}} = 2$ ), according to our definition that



288 tracks must contain exactly three hits, with one hit in each detector layer. This  
 289 follows from the fact that these events are generally produced by the passage of  
 290 a single charged particle, such that the two tracks found in the event are just  
 291 segments of a single charged track.

292 Conversely, the average charged multiplicity from antiproton annihilation  
 293 results in roughly three charged tracks per annihilation. This means that a  
 294 large number of annihilation events (about  $\sim 46\%$  of reconstructed vertices)  
 295 will contain more than two charged tracks ( $N_{\text{tracks}} > 2$ ). However, there is still  
 296 substantial overlap between the signal and background  $N_{\text{tracks}}$  distributions,  
 297 wherein many annihilation events contain only two charged tracks, while some  
 298 cosmic background events are accompanied by particle showers or scattering  
 299 resulting in large numbers of tracks in those events.

#### 300 4.1.2. Combined linear fit residual, $\delta$

301 Cosmic ray particles passing through the ALPHA detector are expected to  
 302 follow, to first order, straight-line trajectories. Thus, events consistent with a  
 303 single, linear, particle track are likely to be the result of the passage of a cosmic  
 304 particle. To test for this case, the hit positions in an event can be fitted with  
 305 a line. The combined linear residual,  $\delta$ , can be used to evaluate how closely an  
 306 event conforms to a single straight line track. This estimator is written as,

$$\delta = \min \left\{ \sum_{i \in \mathcal{F}} d_{\perp,i}^2 \right\}, \quad (3)$$

307 where  $d_{\perp,i}$  is the perpendicular distance, or residual, between the fitted line and  
 308 the  $i$ -th hit in the set of hits,  $\mathcal{F}$ . Finally, the minimization of Eq. 3 involves  
 309 iterating over every pair of tracks. This iteration is done to ensure that, even  
 310 if the event contains several tracks, the combination providing the best fit, or  
 311 smallest value for  $\delta$ , is chosen. The track pair with the smallest  $\delta$  is used to  
 312 characterize the event, as this analysis is explicitly intended to identify cosmic-  
 313 like events.

314 If the hits in an event fit a perfect line,  $\delta$  will evaluate to zero. However,  
 315 due to their curvature in the magnetic field and multiple scattering as they pass

316 through the apparatus and detector, cosmic trajectories often deviate from the  
317 ideal, resulting in a broadening of the  $\delta$  distribution. Annihilation events, on  
318 the other hand, are not expected to produce many co-linear tracks, and should  
319 return values of  $\delta$  well removed from the cosmic distribution. The  $\delta$ -cut is set to  
320 account for the curvature and multiple scattering of cosmic trajectories, while  
321 minimizing the loss of acceptance for annihilation events.

#### 322 4.1.3. Vertex radius, $R$

323 Annihilations must originate from within the trapping region of the appara-  
324 tus, either on the surface of the electrodes, or on background gas. This physical  
325 constraint restricts the possible locations of the reconstructed vertex. In partic-  
326 ular, the radial coordinate of the annihilation vertex is expected to be, within  
327 the radial reconstruction resolution, at the electrode radius. Likewise, it is rea-  
328 sonable to assume that reconstructed vertices far outside the trap volume are  
329 spurious and unphysical.

330 A cosmic event with two co-linear tracks, however, will return a vertex that  
331 is unconstrained in the radial coordinate and often well outside the trapping  
332 volume (an example of such a vertex is Fig. 4 b). Thus, events where the vertex  
333 radius is much greater than the electrode radius are attributed to cosmic rays  
334 and categorized as background.

#### 335 4.2. Datasets

336 Representative sample sets are required for both the annihilation signal and  
337 cosmic background in order to place the cuts on the discriminating variables. In  
338 the case of the ALPHA experiment, the signal and background can be measured  
339 separately, and as such, dedicated data samples can be collected for each.

340 The annihilation signal sample was constructed from 335 cycles where positrons  
341 and antiprotons were mixed together to form antihydrogen in the magnetic field  
342 of the neutral atom trap [9]. Each mixing cycle lasted for 1 s, and a total of  
343 165520 readout events were collected. Over the combined 335 seconds of sig-  
344 nal collection, about 3350 events are expected from cosmic background, which

345 constitutes a contamination of  $\sim 2\%$ .

346 Conversely, the background sample set was collected by operating the detec-  
347 tor with no antiparticles present within the apparatus. So as to best emulate the  
348 situation of interest, the neutral trap magnets were kept engaged throughout  
349 the background collection. Data were recorded over several periods, which to-  
350 taled almost 3 hours, with 109824 readout events. With no antiparticles present,  
351 these events should all be background signals, from such sources as cosmics or  
352 electronic noise-induced detections.

### 353 *4.3. Cut placement and optimization*

354 The separation of signal from background is only effective with well placed  
355 cuts on the discriminating variables. These cuts are determined so as to maxi-  
356 mize the number of cosmic events rejected, while retaining as many annihilation  
357 events as possible. This analysis is focused on optimizing the expected signal  
358 significance during attempts to trap antihydrogen. An appropriate figure of  
359 merit for this optimization is the p-value for a Poisson distribution,  $\alpha$ . This  
360 p-value represents the probability of observing  $n_0$  events (or more) solely due  
361 to a fluctuation in a Poisson-distributed background with mean  $b$ . Here, the  
362 expected signal and expected background can be parametrized as functions of  
363 cuts on the combined linear residual,  $\delta_{\text{cut}}$ , and vertex radius,  $R_{\text{cut}}$ , as well as  
364 categorized according to the number of charged tracks,  $N_{\text{tracks}}$ , found in the  
365 event. The p-value figure of merit can then be written as

$$\alpha(n_0, b) = \sum_{n=n_0}^{\infty} \frac{b^n e^{-b}}{n!}, \quad (4)$$

366 where the observed events and background rate depend on the cuts used ( $n_0 =$   
367  $n_0(R_{\text{cut}}, \delta_{\text{cut}}, N_{\text{tracks}})$  and  $b = b(R_{\text{cut}}, \delta_{\text{cut}}, N_{\text{tracks}})$ ). The signal optimization  
368 then proceeds by finding the set of cuts which minimize  $\alpha(n_0, b)$ , or equiva-  
369 lently, maximize the signal significance. In order to avoid unintentional bias,  
370 all analyses were performed and finalized on the auxiliary datasets described in  
371 Sec. 4.2. Moreover, the cut optimization was performed blindly, i.e. without  
372 direct reference to the trapping data.

373 The events are first separated into two categories: those with two charged  
 374 tracks ( $N_{\text{tracks}} = 2$ ), and events with more than two charged tracks ( $N_{\text{tracks}} >$   
 375 2). This categorization separates the events into background-dominated, and  
 376 signal-dominated sets, as the vast majority of background events fall into the  
 377  $N_{\text{tracks}} = 2$  category (see Fig. 4(b) for an example of a typical background  
 378 event). The  $N_{\text{tracks}} > 2$  set then contains mostly signal events, with some back-  
 379 ground events to be rejected, while the  $N_{\text{tracks}} = 2$  category contains primarily  
 380 background events, with some signal events to be extracted. Thus, cuts on the  
 381  $N_{\text{tracks}} = 2$  events should be relatively stringent to reject as many background  
 382 events as possible, while the  $N_{\text{tracks}} > 2$  cuts should be much more inclusive.

383 In order to optimize the expected signal significance, Eq. 4 is estimated over  
 384 a wide range of radius and residual cuts, as well as for both  $N_{\text{tracks}}$  categories.  
 385 Since  $b$  and  $n_0$  are functions of the applied cuts, both parameters must be de-  
 386 termined for every set of cuts. The background rate can easily be determined  
 387 by directly applying the cuts to the cosmic background dataset and examin-  
 388 ing the surviving distribution. This method has the advantage of an accurate  
 389 cosmic background estimate, since a direct measurement can be made. How-  
 390 ever, it is difficult to estimate the expected signal rate, as the dynamics of the  
 391 antihydrogen distribution are not well characterized throughout the trapping  
 392 experiments. Instead, a baseline number of signal events is taken from Ref. [28]  
 393 and the expected number of events is determined for each set of cuts by scaling  
 394 the baseline value according to the auxiliary annihilation distribution.

395 For a low-rate process such as antihydrogen trapping,  $n_0$  is assumed to fol-  
 396 low Poisson statistics. To reflect these statistics, an aggregate value for  $\alpha$  is  
 397 calculated (for each set of cuts) using 5000 pseudo-experiments. Each pseudo-  
 398 experiment is performed as follows:

- 399 1. A Poisson distribution with mean  $n_0$  is sampled so as to obtain an pseudo-  
 400 experimental number of observed events,  $n_s$ .
- 401 2. The p-value for the pseudo-experiment is calculated  $\alpha_i = \sum_{n=n_s}^{\infty} b^n \exp(-b)/n$ .
- 402 3. The representative value for the ensemble of p-values is taken as the

$N_{\text{tracks}}$	Vertex radius, $R_{\text{cut}}$ (cm)	Linear residual, $\delta_{\text{cut}}$ (cm <sup>2</sup> )
= 2	< 4	> 2
> 2	< 4	> 0.05

Table 2: Final parameter cut conditions. Events satisfying these conditions are classified as annihilations.

403 log-average  $\alpha = \exp(1/n \sum_{i=1}^n \ln \alpha_i)$ . This measure of central tendency  
404 takes into account the logarithmic nature of the distribution of pseudo-  
405 experimental p-values.

406 Fig. 6 shows the parameter space of the figure of merit as a function of the  
407 discriminating variables. The p-value is expressed in terms of standard devia-  
408 tions for a one-sided normal distribution, such that maximizing the expected  
409 significance corresponds to minimizing the expected p-value. These distribu-  
410 tions are constructed by considering an array of  $(R_{\text{cut}}, \delta_{\text{cut}})$  cut value pairs.  
411 The final set of cuts were then chosen so as to take into account the expected  
412 significance, and the final choices of cuts are shown as the black crosses in Fig.  
413 6, and enumerated in Table 2. The optimal placement of the cut values is un-  
414 affected when the estimated signal rate is varied, indicating that this analysis  
415 provides a robust determination of background rejection cuts.

#### 416 4.4. Results of the background rejection

417 The effectiveness of the cut placement optimization can be evaluated by  
418 applying these cuts to the datasets described in Section 4.2. Fig. 7 shows  
419 the distributions of the discriminating variables for both signal and background  
420 datasets, along with the distributions after the application of the cuts. By apply-  
421 ing the cuts to the background sample,  $(99.54 \pm 0.02)\%$  of the events are rejected,  
422 corresponding to a background acceptance rate of  $(47 \pm 2) \times 10^{-3}$  events/s.

423 Similarly,  $(75.7 \pm 0.1)\%$  of events in the signal sample return a vertex and  
424  $(85.1 \pm 0.1)\%$  of those vertices survive the cuts, which combined give a signal  
425 acceptance of  $(64.4 \pm 0.1)\%$ . The overall detection efficiency is then  $(58 \pm$   
426  $7)\%$ , which is the product of the trigger efficiency  $(90 \pm 10)\%$  and the signal

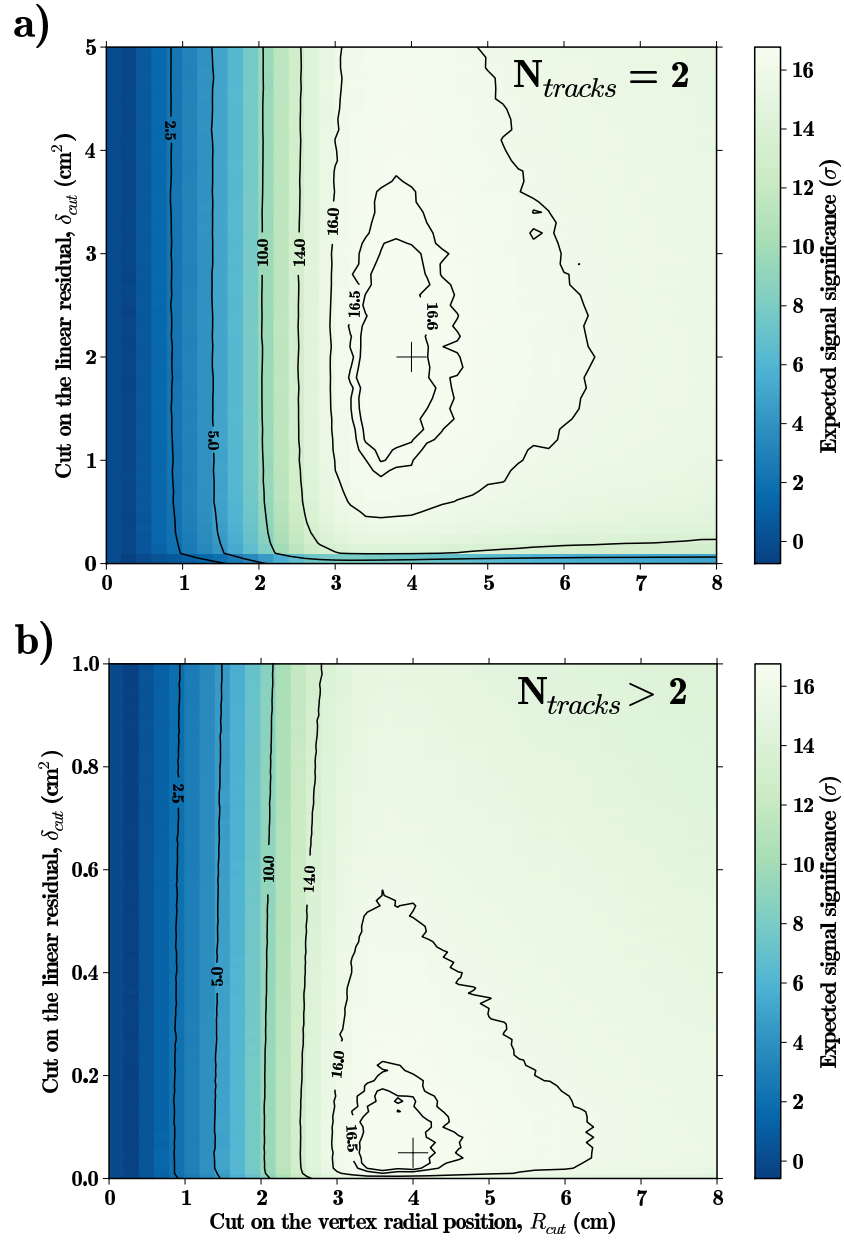


Figure 6: Contour plots for the expected signal significance as a function of the cuts on the vertex radius,  $R_{cut}$ , and combined linear residual,  $\delta_{cut}$ , for a) events with  $N_{tracks} = 2$ , and b) events with  $N_{tracks} > 2$ . The final cut decisions are shown as the black crosses.

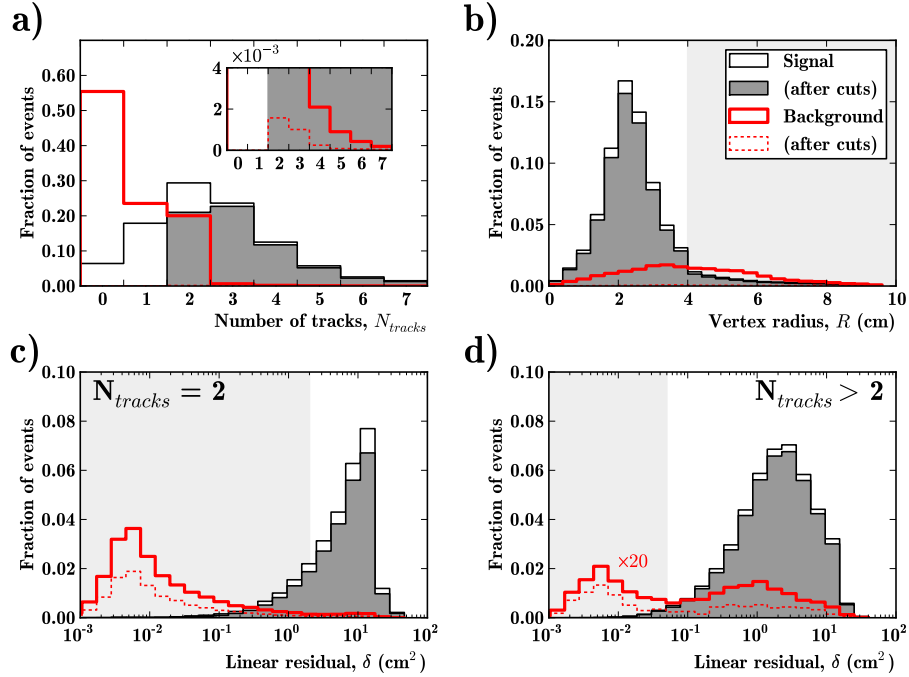


Figure 7: Measured signal and background distributions for the discriminating variables. Shown are (a) the distribution of the number of charged tracks per event,  $N_{tracks}$ , (b) the radial component of the reconstructed vertex,  $R$ , and the combined linear residuals for the cases where (c)  $N_{tracks} = 2$ , and (d)  $N_{tracks} > 2$ . The annihilation signal sample is shown as the solid black trace, while the signal sample after the application of all other cuts than the one plotted is shown as the solid grey filled trace. The cosmic background sample is shown as the solid thick trace, and the dashed trace shows the background sample after the application of all other cuts than the one plotted. All distributions are normalized to the number of events in each sample set. The inset in (a) shows a magnification of the y-axis, highlighting the effect of the cuts on the background sample. Likewise, both background traces have been multiplied by a factor of 20 in (d) to allow for comparison. The lightly shaded areas represent the regions rejected by the cuts.

427 acceptance. Since the events included in these datasets were collected from *in*  
428 *situ* measurements (Sec. 4.2), the charge-collection and threshold efficiencies of  
429 the silicon modules are convolved into the above numbers.

## 430 5. Application to trapped antihydrogen detection

431 A strong motivation for the inclusion of a silicon detector in the ALPHA  
432 experiment is that it is sensitive to the annihilations of individual antiprotons.  
433 This is especially relevant for antihydrogen trapping experiments, where the  
434 expected number of trapped atoms is very small. During the 2009 experimental  
435 beamtime, 212 trapping experiments were completed, combining  $10^7$  antiprotons  
436 with  $1.3 \times 10^9$  positrons [28]. Each trapping experiment involved the mixing of  
437 positrons and antiprotons to synthesize antihydrogen within the neutral atom  
438 trap fields, such that sufficiently low-energy antihydrogen atoms would be unable  
439 to escape the magnetic trap. Then, shortly after mixing was stopped, the neutral  
440 trap was quickly de-energized, allowing any trapped antihydrogen atoms to  
441 escape and annihilate. In total, 36 detector readout events were recorded in  
442 the 30 ms window during the fast shutdown of the neutral trap. After a blind  
443 analysis to determine the optimal cut placement (following Sec. 4), 6 events  
444 were identified that satisfied all selection criteria [28]. The cosmic ray event  
445 suppression provided a background acceptance rate of  $(2.2 \pm 0.1) \times 10^{-2}$  events/s  
446 during these attempts (due to the specifics of the detector trigger used), the  
447 probability that all 6 events observed were due to statistical fluctuations in the  
448 cosmic ray background is  $9.2 \times 10^{-9}$ , corresponding to a signal significance of  
449  $5.6\sigma$ .

450 In addition to the cosmic ray particles, bare antiprotons that have been  
451 mirror-confined in the inhomogeneous magnetic field of the neutral-atom trap  
452 are another possible background [28]. Mirror-trapped antiprotons are a difficult  
453 background to isolate, as the annihilation signature of the bare antiproton is  
454 identical to that of released antihydrogen. However, mirror-trapped antipro-  
455 tons can be ruled out as a source of annihilation signal by applying a static



456 electric field during the shutdown of the magnetic trap. This bias field acts to  
457 axially deflect any charged particles, while leaving the neutral atoms unaffected  
458 [1]. This method of discriminating between released antihydrogen atoms and  
459 bare mirror-trapped antiprotons relies crucially on the axial reconstruction of  
460 the annihilation position. Specifically, while the release of trapped antihydrogen  
461 results in a vertex distribution that is axially extended across the length (14 cm)  
462 of the trap, mirror-trapped antiprotons will annihilate in an axially narrow dis-  
463 tribution [29]. Moreover, the applied bias field can be used to positively identify  
464 the presence of mirror-trapped antiprotons by shifting the average axial vertex  
465 position by as much as 14 cm, while leaving the distribution of antihydrogen ver-  
466 tices unchanged. The differences between these vertex distributions are readily  
467 observable, given the axial reconstruction resolution of 0.56 cm FWHM stated  
468 in Sec. 3.2.

469 During the 2010 AD beamtime, ALPHA performed 335 trapping experi-  
470 ments (similar to the experiments described above), combining a total of  $10^7$   
471 antiprotons with  $7 \times 10^8$  positrons. Crucially, almost a third of those trapping  
472 experiments involved the bias field to deflect any mirror-trapped antiprotons.  
473 Overall, 307 detector readout events were recorded during the 30 ms detec-  
474 tion window, and 38 events satisfied all of the annihilation selection criteria [1].  
475 Moreover, the selected events formed an axial vertex distribution characteris-  
476 tic of released antihydrogen atoms, and inconsistent with bare mirror-trapped  
477 antiprotons. The confinement time during the initial trapping experiments was  
478 set to 172 ms, which was the shortest time required to perform the measure-  
479 ment. However, following refinements to the trapping procedure, the number  
480 of trapped atoms per attempt was increased by up to a factor of five and con-  
481 finement times were lengthened to as long as 1000 s [2]. In total, 309 trapped  
482 antihydrogen annihilation events were recorded and examined during the 2010  
483 experiments.

## 484 **6. Summary**

485 The ALPHA collaboration has constructed a silicon annihilation reconstruction  
486 detector for the purposes of detecting and studying antihydrogen. This article  
487 describes the methods related to the reconstruction of the vertex position  
488 of an antiproton annihilation. In addition, the analysis to optimize the back-  
489 ground suppression is presented. After optimization, these algorithms permit  
490 a background rate of  $(47 \pm 2) \times 10^{-3}$  events/s in the ALPHA detector, while  
491 accepting  $(64.4 \pm 0.1)\%$  of the recorded annihilation events. The detector and  
492 methods described above were crucial to the successful observation of trapped  
493 antihydrogen, and will likely be an important part of future spectroscopic mea-  
494 surements in ALPHA.

## 495 **7. Acknowledgements**

496 This work was supported by CNPq, FINEP/RENAFAE (Brazil); ISF (Is-  
497 rael); MEXT (Japan); FNU (Denmark); VR (Sweden); NSERC, NRC/TRIUMF,  
498 AIF/AITF, FQRNT, and the Killam Trust (Canada); the DOE and the NSF  
499 (USA); and EPSRC, the Royal Society and the Leverhulme Trust (UK). We  
500 would like to thank D. Seddon, J. Thornhill, and D. Wells (University of Liver-  
501 pool) for their work on the construction of the vertex detector.

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